Impact of noise sources and optical design on defect detection sensitivity in extreme ultraviolet actinic pattern inspection tool

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Abstract. We discuss the impact of various noise sources and the optical design in bright field extreme ultraviolet (EUV) actinic inspection of mask features for defects in the patterned absorber. It is shown that an optimum pixel size is needed to maximize the defect signal-to-noise ratio (SNR) to balance the trade-off in increasing signal strength with shot noise from defect signal and the background pattern intensity (mask layout image) and speckle noise from the mask blank roughness. Moreover, we consider defocus showing that the EUV mask phase effect has an asymmetric impact on pattern defect SNR's through-focus behavior. The impact of defocus limits inspection performance based on defect SNR. Using critical defect sizes in a case study, we show the defect SNR performance of the limiting case and discuss the possibility of utilizing a nominal defocus in the inspection system to leverage the phase effect of EUV mask absorber to improve the defect SNR. A 50% improvement in defect SNR is shown to be possible by introducing a −50 nm nominal defocus into the bright field inspection system.

Keywords: extreme ultraviolet actinic pattern inspection; extreme ultraviolet mask pattern defect; extreme ultraviolet lithography; optical design; signal-to-noise ratio.

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1 Introduction

As extreme ultraviolet (EUV) lithography marches toward high-volume manufacturing, the resolution limitations of deep ultraviolet inspection and the introduction of EUV pellicle make EUV actinic pattern inspection a key factor in the mask-making process flow. At-wavelength (actinic) inspection provides a better optical resolution on smaller defects for advanced technology nodes and enables through-pellicle inspection. Therefore, actinic inspection tools with high sensitivity and large throughput will be highly valuable to the inspection. Therefore, actinic inspection tools with high sensitivity and large throughput will be highly valuable to the mask-making process flow. At-wavelength (actinic) inspection provides a better optical resolution on smaller defects for advanced technology nodes and enables through-pellicle inspection. Therefore, actinic inspection tools with high sensitivity and large throughput will be highly valuable to the EUV mask fabrication process.

Here, we study the relationship between various noise sources and the optical design. Noise sources in pattern mask inspection include speckle noise from the surface roughness, camera noise from the dark current of the charged-coupled device (CCD) camera, and photon shot noise from the background pattern and defect signal intensities. Additionally, pixel size and photon level (source power) are critical factors in determining the throughput and defect sensitivity. With a fixed number of photons per pixel, a larger pixel size might increase the throughput but lower both the defect signal strength and noise. On the other hand, with a fixed number of photons/nm², a larger photon count per pixel achieved by increasing the pixel size might increase the pixel signal strength but also increase the photon shot noise from the defect signal and background pattern intensity.

To gain a better understanding of these tradeoffs and interplays, in this manuscript we first discuss the impact of each noise source and defocus under various pixel sizes and photon densities on defect signal-to-noise ratio (SNR). Then, we calculate the defect sensitivity for a set of critical defects and identify the limiting case, which has the smallest defect SNR. In the final section, we discuss the impact of defocus and the possibility of introducing a nominal defocus into the inspection system to operate at a higher defect SNR region to achieve a better defect capture rate.

2 Background

2.1 Simulation Parameters and Defect Signal-to-Noise Ratio Definitions

The simulation study presented here uses a thin mask 2-D model to generate the EUV (13.5 nm) images. Details on the determination of the 2-D mask file with all the necessary information for the defect SNR calculation is described in Sec. 2.2. For the imaging conditions, we assume a bright field imaging mode with a numerical aperture (NA) of 0.16 and a disk illumination with a sigma value of 0.3. For the detector conditions, we assume a pixel size ranging from 10 to 50 nm in mask scale, and a photon level of 1000 photons/pixel or 10 to 50 photons/nm² depending on the pixel size. The patterns used in the simulation are dense line and isoline perpendicular to the 6-deg incident illumination angle with 64-nm halfpitch in mask scale, and also dense contact pattern with 80-nm half pitch in mask scale. Square-shaped extrusion and intrusion defects with size ranging from 6.4 × 6.4 nm² to 40 × 40 nm² are used. The definition of extrusion/intrusion defect and the optical properties of the absorber materials used for the EUV mask...
pattern are shown in Fig. 1. For speckle, we assume a root-mean-square (RMS) mask roughness of 61 pm with a correlation length of $\sim 100$ nm. In the analysis, we adopt a die-to-database defect detection approach, subtracting the ideal (noise free) background pattern image. Equation (1) shows the defect SNR definition used in the following discussion.\(^1\) Photon shot noise includes the impact from both defect signal and pattern background intensity:

$$\text{SNR} = \frac{\text{Defect signal}}{\text{Speckle noise} + \text{camera noise} + \text{photon shot noise}}. \quad (1)$$

2.2 Defect Signal-to-Noise Ratio Process Flow

To include both the absorber and substrate properties into the pattern mask inspection modeling, we extract the patterned mask electric field from a 3-D model\(^2\) and mix it with the 2-D mask roughness electric field, as shown in Fig. 2. It has previously been shown that the multilayer roughness effect is adequately modeled with the single surface approximation.\(^3\) This hybrid approach allows us to include not just the mask 3-D effect of EUV mask, but also the interaction between mask roughness and pattern electric field modulation while keeping reasonable constraints on the size of the computation.

From the densely sampled aerial image, we utilize pixel binning to mimic the results under different pixel sizes for both die and database images. Also, in our modeling of the inspection process, we assume the effective signal to be derived based on a $2 \times 2$ pixel convolution of the image. Figure 2 shows the aerial image as we change the pixel size from 10 to 30 nm. In the final step, we include system noise, such as photon noise and camera noise, to calculate the defect SNR from die-to-database images, as those shown in Fig. 2.

The goal of actinic inspection of the patterned absorber is to identify locations, where significant deviations from design in the absorber shape occur and to which repair should be applied.

2.3 Impact of Defocus on Defect Signal-to-Noise Ratio Through-Focus Behavior

In previous studies,\(^4,5\) it was shown that the impact of the pattern phase effect due to the phase-shifting of absorber materials causes the pattern defect to exhibit a mixed (phase + absorber) behavior. The impact of this, as shown in Fig. 3, is that the peak defect signal is no longer at the best focus position as an ideal absorber defect. Moreover, the interaction between mask roughness and pattern mask also causes an asymmetric through-focus behavior for speckle...
noise as shown in Fig. 3, and thus an asymmetric through-focus SNR.

3 Simulation Results

3.1 Defect Signal Versus Speckle Noise Under Various Pixel Sizes and Defocus Positions

In the first step, we only consider the defect signal and speckle noise in a defect SNR calculation in order to understand the impact of pixel size on both terms. The example defect used here is a dense line intrusion defect with a size of $26 \times 26$ nm$^2$. As shown in Fig. 4(a), the normalized signal strength and speckle noise drop as the pixel size increases. Moreover, Fig. 4(b) shows that signal drops faster than speckle noise as the pixel size increases causing smaller defect SNR when we use larger pixel sizes in the inspection tool.

The drastic change of defect SNR, when the pixel size is larger than 25 nm, is due to the pixel area exceeding the optical resolution. The optical resolution limit for our system is 51 nm based on the NA (0.16) and the wavelength (13.5 nm). Therefore, for the sample defect we used here, which is below the resolution limit, the FWHM of its defect intensity in the aerial image is about 50 nm. Also, in our modeling of the inspection process, we assume the effective signal to be derived based on a $2 \times 2$ pixel convolution of the raw camera image raising the effective inspection pixel size to 50 nm. When the camera pixel size is smaller than 25 nm, the defect intensity distribution is larger than the effective pixel size. Thus, increasing the pixel size also leads to increased peak signal. Once the pixel size becomes larger than 25 nm, the convolved area is larger than the defect intensity distribution and the area without the defect signal lowers the defect signal.

As shown in Fig. 4(c), the impact of the phase associated with the pattern defect causes asymmetric defect SNR through focus. The best focus position used in this paper is defined as the position that has minimum speckle contrast, as shown in Fig. 3. Defocus position $= +100$ nm (image closer to the lens) has a smaller defect SNR compared with the other two cases of best focus and defocus of $-100$ nm. The asymmetric behavior is the limiting factor on the defect SNR performance if we operate the inspection tool near the best focus position.

3.2 Defect Signal Versus Speckle Noise and Camera Noise Under Various Pixel Sizes and Defocus Positions

For system noise, dark current in the CCD camera is the primary noise source that impacts the defect SNR. Therefore, we discuss the impact of camera noise, which is assumed to be a constant with varying pixel sizes, on the defect SNR trend under different pixel sizes in this section. As shown in Fig. 5(a), the defect SNR at various pixel sizes drops significantly and the trend is different from the previous case with the introduction of another noise source into the defect signal distribution. As shown in Fig. 4(c), the impact of the phase associated with the pattern defect causes asymmetric defect SNR through focus. The best focus position used in this paper is defined as the position with minimum speckle contrast, as shown in Fig. 3.

Fig. 4 (a) Normalized defect signal and speckle noise under various pixel sizes relative to the value at 10-nm pixel size. (b) Defect SNR under various pixel sizes, only defect signal and speckle noise are taken into the defect SNR calculation. (c) Defect SNR under various pixel sizes and different defocus positions: Defocus position $= -100$, 0, and $+100$ nm. Defect type: dense line intrusion defect with a size about $26 \times 26$ nm$^2$ on the mask. The best focus position is defined as the position with minimum speckle contrast, as shown in Fig. 3.
SNR calculation. The reason is that the speckle noise caused by the phase-dominated mask surface roughness at focus under our illumination and optic settings (bright field) is much smaller than the camera noise. The RMS of the camera noise is 12 (unit: √Photons) while the speckle noise at focus is only 5.5 (unit: √Photons). Therefore, the noise term in Eq. (1) is dominated by a constant camera noise and the overall trend of defect SNR is determined by the defect signal: smaller signal strength as pixel size increases. Figure 5(b) shows the result at various defocus positions. The results from the best focus and defocus positions =−100 nm have similar defect SNR trends while the defocus position = +100 nm has a smaller defect SNR performance. The best focus position is again defined as the position that has minimum speckle contrast, as shown in Fig. 3.

### 3.3 Defect Signal Versus Speckle Noise, Camera Noise, and Photon Shot Noise Under Various Pixel Sizes and Defocus Positions

Next, we consider the impact of photon shot noise, originating from the defect signal and background pattern intensity. With fixed photon densities (source power), larger pixel size means more photons in a single pixel. For defect signal strength, it means that the signal strength in the number of photons is larger simply by virtue of increasing the pixel size. However, the corresponding photon shot noise (√Photons) from both defect signal and background pattern intensity increases as well. The interaction between defect SNR and pixel size under fixed photon density is shown in Fig. 6(a).

There exists an optimum pixel size to maximize the defect SNR under fixed photon density. This is due to the fact that the increasing defect signal strength is compensated by the increase of photon shot noise. Moreover, under fixed pixel size, the defect SNR improvement is smaller as you increase the photon density. When the pixel size is 25 ×25 nm², the defect SNR improves 25% with a 2× increase of photons per pixel from 10 to 20 photons/nm². However, the defect SNR only improves 8% with a 1.5× increase of photons per pixel from 20 to 30 photons/nm².

Figure 6(b) shows the result at various defocus positions when we include all the noise terms in the SNR calculation as shown in Eq. (1). Under fixed photon density, there is an optimum pixel size to reach the highest defect SNR. Moreover, due to the asymmetric impact from defocus, defocus position =−100 nm and the best focus position have similar defect SNR performances and defocus position =+100 nm has the smallest defect SNR.

### 3.4 Critical Defect Case Study and How to Improve the Defect Signal-to-Noise Ratio for Limiting Case

Based on the results shown in the previous sections, we can calculate the defect SNR performance for defects that cause 10% CD variation on the pattern aerial image. To determine the critical defects for dense line, isoline, and dense contact patterns, we assume projection lithography tool imaging parameters as opposed to the inspection mode parameter we described in Sec. 1. The projection lithography tool imaging
parameters we assume include an NA of 0.33 and dipole illumination with a sigma value between 0.2 and 0.9 with 90-deg opening angle for dense line pattern, disk illumination with a sigma value 0.5 for isoline pattern, and quasar with a sigma value between 0.2 and 0.9 with 45-deg opening angle for contact pattern. Table 1 shows the critical defect of each situation. The smallest critical defect is the extrusion defect for the isoline pattern, which is about $13 \times 13 \text{ nm}^2$ on the mask. Figure 7 shows the critical defects SNR results under various pixel sizes and photon densities with defocus position $= +100 \text{ nm}$, which was the worst case studied above. As shown in Fig. 7(b), the extrusion defect SNR is in the range of 3 to 7 no matter how we vary the pixel size and photon density. This defect would be the limiting case among all critical defects since it has the smallest defect SNR and thus capture rate.

Figure 8 shows the isoline extrusion critical defect through-focus SNR result, with pixel size and photon density of $30 \times 30 \text{ nm}^2$ and 50 photons/\text{nm}^2 respectively. A precise defocus range control is needed to reach the desired defect SNR since the defect SNR is sensitive to the defocus position as shown in Fig. 8. Moreover, the defect signal and speckle

<table>
<thead>
<tr>
<th>Pattern type</th>
<th>Defect type</th>
<th>Size (nm$^2$)</th>
</tr>
</thead>
<tbody>
<tr>
<td>Dense line</td>
<td>Intrusion</td>
<td>$25.6 \times 25.6$</td>
</tr>
<tr>
<td></td>
<td>Extrusion</td>
<td>$19.2 \times 19.2$</td>
</tr>
<tr>
<td>Isoline</td>
<td>Intrusion</td>
<td>$19.2 \times 19.2$</td>
</tr>
<tr>
<td></td>
<td>Extrusion</td>
<td>$12.8 \times 12.8$</td>
</tr>
<tr>
<td>Dense contact</td>
<td>Intrusion</td>
<td>$16 \times 16$</td>
</tr>
</tbody>
</table>

Table 1 Critical defect size for different pattern designs.

Fig. 7 Critical defect SNR at defocus position $= +100 \text{ nm}$ for various patterns: (a) dense line, (b) isoline, and (c) dense contact. Legends in (c) indicate the photon density level for each curve, ranging from 10 to 50 photons/\text{nm}^2. The best focus position is defined as the position with minimum speckle contrast, as shown in Fig. 3.
noise both have an asymmetric through-focus behavior due to the phase effect of EUV mask. For the isoline extrusion critical defect, SNR$_{\text{min}} = 6.8$ at defocus position $= +100$ nm. However, if we introduce a nominal defocus of $-50$ nm and operate the inspection tool in a defocus range: $-150$ to $+50$ nm, we can improve the SNR$_{\text{min}}$ by operating the tool at a higher defect SNR region. With this new setting, SNR$_{\text{min}} = 10.3$ at defocus position $= +50$ nm, a 50% improvement in SNR. Moreover, it is interesting to note that the new best focus position (defocus $= -50$ nm) of the inspection tool corresponds to neither the maximum aerial image contrast nor minimum speckle contrast.

4 Conclusion

In this paper, we discuss the impact of various noise sources and the optical design on defect SNR in EUV actinic pattern mask inspection. It is found that signal drops faster than speckle noise as pixel size increases due to the resolution limit of the defect aerial image intensity distribution. We also show that even though larger pixels can lead to more photons per pixel under fixed photon density (source brightness), the defect SNR is smaller due to the increase of both defect signal and photon shot noise ($\sqrt{\text{Photons}}$) from the defect signal and the background pattern intensity. The improvement of defect SNR by increasing photon density for a fixed pixel size also saturates for the same reason. The asymmetric impact of defocus reduces the defect SNR$_{\text{min}}$ when operating at defocus position $> 0$. In the critical defect case study, we show that the smallest critical defect has a possible but narrow defocus range to operate at high defect SNR (SNR $\geq 10$), and the introduction of a nominal defocus into the inspection system can utilize this defocus range with higher defect sensitivity. A 50% improvement on SNR is achieved by introducing a $-50$ nm nominal defocus into the inspection system.

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References


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